

Carbon under extreme conditions: Phase boundaries and electronic properties from first-principles theory

Alfredo A. Correa^{*†}, Stanimir A. Bonev^{†‡}, and Giulia Galli^{*§}

^{*}Department of Physics, University of California, Berkeley, CA 94720; [†]Lawrence Livermore National Laboratory, Livermore, CA 94550; and [‡]Department of Physics, Dalhousie University, Halifax, NS, Canada B3H 3J5

Communicated by Russell J. Hemley, Carnegie Institution of Washington, Washington, DC, December 8, 2005 (received for review October 6, 2005)

At high pressure and temperature, the phase diagram of elemental carbon is poorly known. We present predictions of diamond and BC8 melting lines and their phase boundary in the solid phase, as obtained from first-principles calculations. Maxima are found in both melting lines, with a triple point located at ≈ 850 GPa and $\approx 7,400$ K. Our results show that hot, compressed diamond is a semiconductor that undergoes metalization upon melting. In contrast, in the stability range of BC8, an insulator to metal transition is likely to occur in the solid phase. Close to the diamond/liquid and BC8/liquid boundaries, molten carbon is a low-coordinated metal retaining some covalent character in its bonding up to extreme pressures. Our results provide constraints on the carbon equation of state, which is of critical importance for devising models of Neptune, Uranus, and white dwarf stars, as well as of extrasolar carbon-rich planets.

phase transitions | melting | high pressure | molecular dynamics | metalization

Elemental carbon has been known since prehistory, and diamond is thought to have been first mined in India $>2,000$ years ago, although recent archaeological discoveries point at the possible existence of utensils made of diamond in China as early as 4,000 before Christ (1). Therefore, the properties of diamond and its practical and technological applications have been extensively investigated for many centuries. In the last few decades, after the seminal work of Bundy and coworkers (2) in the 1950s and '60s, widespread attention has been devoted to studying diamond under pressure (3). For example, the properties of diamond and, in general, of carbon under extreme pressure and temperature conditions are needed to devise models of outer planet interiors (e.g., Neptune and Uranus) (4–6), white dwarfs (7, 8) and extrasolar carbon planets (9).

Nevertheless, under extreme conditions the phase boundaries and melting properties of elemental carbon are poorly known, and its electronic properties are not well understood. Experimental data are scarce because of difficulties in reaching megabar (1 bar = 100 kPa) pressures and thousands of Kelvin regimes in the laboratory. Theoretically, sophisticated and accurate models of chemical bonding transformations under pressure are needed to describe phase boundaries. In most cases, such models cannot be simply derived from fits to existing experimental data, and one needs to resort to first-principles calculations, which may be very demanding from a computational standpoint.

It has long been known that diamond is the stable phase of carbon at pressures above several gigapascal (2). Total energy calculations (10, 11) based on Density Functional Theory (DFT) predict a transition to another fourfold coordinated phase with the BC8 symmetry[¶] at $\approx 1,100$ GPa and 0 K, followed by a transition to a simple cubic phase at pressures $>3,000$ GPa. These transitions have not yet been investigated experimentally, because the maximum pressure reached so far in diamond anvil cell experiments on carbon is 140 GPa (14). The only experimental information on diamond melting in the approximately 10^2 - to 10^3 -GPa range is from recent shock-wave experiments, where melting and a transition to a conducting fluid were observed (15). However, these measure-

ments were insufficient to locate phase boundaries and inconclusive as to whether diamond undergoes an insulator-to-metal transition before or after melting.

A number of experimental studies (16, 17) and *ab initio* simulations (18–20) have focused on the phase boundaries close to the graphite–diamond–liquid triple point ($\approx 4,000$ K and 15 GPa). Unlike other solids with the diamond structure (e.g., Si and Ge), at these pressures the melting line of carbon has a positive slope (19). However, at much higher pressure (several hundred GPa), Martin and Grumbach (21) suggested that the slope becomes negative. This result was inferred from the comparison of the specific volumes of liquid and solid carbon computed from *ab initio* molecular dynamics. Empirical potentials (22–25) have also been used to investigate melting of carbon with different degrees of success. In general, empirical potentials lack predictive power over large density variations and, more specifically, important discrepancies between *ab initio* calculations and empirical potentials have been recently found, e.g., for the existence of a liquid/liquid phase transition at low pressure (20, 24, 26).

In this paper, we report on the properties of carbon in the pressure and temperature range of 15–2,000 GPa and 0–10,000 K, as obtained from first-principles calculations based on DFT. We have investigated solid/liquid phase boundaries and analyzed structural and electronic transformations as pressure and temperature are increased. In particular, we have determined diamond and BC8 melting lines and predicted the location of the diamond/BC8/liquid triple point, inferred from the crossing of computed phase boundaries. Structural and electronic properties of diamond and BC8, and the solid/solid transition in this range of temperatures have also been investigated by a combination of different techniques.

Results

We first computed the equation of state of diamond at zero temperature to test the accuracy of our theoretical and numerical approach. The agreement with the available experimental data in the range 0–140 GPa (14) is excellent. For example, the computed equilibrium volume only differs by 1% from the experimental value, and optical phonon frequencies agree with experiment within 3% (14). This level of accuracy is already obtained with 64-atom supercells and a Γ -point sampling of the Brillouin zone, as evident from convergence tests performed with supercells containing up to 216 atoms and with unit cells with up to $12 \times 12 \times 12$ *k*-point meshes. Consistent with previous studies, we find that at zero temperature the transition pressure from diamond to BC8 is 1,075 GPa. Other candidate structures previously proposed in the liter-

Conflict of interest statement: No conflicts declared.

Freely available online through the PNAS open access option.

Abbreviations: DFT, Density Functional Theory; GGA, generalized gradient-corrected approximation.

[§]To whom correspondence should be sent at the present address: Department of Chemistry, University of California, Davis, CA 95616. E-mail: gagalli@ucdavis.edu.

[¶]BC8 has been experimentally shown to be a metastable phase of silicon (12) and germanium (13) and therefore proposed as a possible high-pressure phase of carbon.

© 2006 by The National Academy of Sciences of the USA

ature, such as R8 (27), ST12 (28), and hexagonal diamond (29), were also examined but were found to be unstable relative to BC8 in the pressure range studied here.^{||}

Calculations of diamond and BC8 melting lines were carried out by direct simulation of solid/liquid coexistence (i.e., by using a so-called two-phase simulation method). This approach has been recently shown to be applicable and efficient within the framework of *ab initio* molecular dynamics simulations of comparable scale (30–32). The two-phase method consists of choosing a given thermodynamic condition in (P , T) space, where liquid and solid systems are prepared and equilibrated. The phases are then put in contact and constant pressure–constant temperature simulations are carried out to determine which of the two phases is more stable at that given (P , T). As simulation progresses, the more stable phase eventually fills the entire volume available in the simulation. The melting line is located by repeating this process at different (P_n , T_n) points. Several complementing criteria, including diffusion coefficients, were used to determine whether melting or crystallization occurs. In addition, correlation functions, coordination numbers, and local order parameters (33) were computed throughout all simulations to monitor the time evolution of the solid/liquid interfaces.

In the simulations of diamond/liquid coexistence and BC8/liquid coexistence, we used 128- and 256-atom cells, respectively, with half of the atoms initially in a solid state and the other half in a liquid state. For selected thermodynamic conditions, we also repeated calculations of diamond/liquid coexistence with 256-atom cells (i.e., doubled in the direction perpendicular to the solid–liquid interface) to assess errors introduced by finite size effects on the calculated melting lines. In addition, we compared two-phase simulations carried out with finite and zero electronic temperature for selected points in phase space; at pressures $>1,000$ GPa, we assessed the error of the pseudopotential approximation by repeating some of our simulations using a pseudopotential with a smaller cut-off radius (decreased from 1.4 to 1.14 a.u. and, consequently, a higher plane wave cutoff, increased to 70 Ryd). We used s , p nonlocal, and d local potentials. In contrast to the low-pressure regime, in which the bonding is dominated by s – p hybridization, at extreme pressures, an accurate description of the d channel is important because the coordination of the liquid becomes >4 and configurations other than tetrahedral are present in the molten phase.

Finally, we determined the diamond/BC8 phase line using two independent approaches: (i) by combining results from two-phase simulations with equation of state calculations near the diamond/BC8/liquid triple point and (ii) from calculations of the free energies of diamond and BC8 based on the quasiharmonic approximation and including quantum ionic motion.

Discussion

Phase Boundaries. The computed melting lines of diamond and BC8 are shown in Fig. 1. The internal consistency of our simulations was verified by computing the slope of melting lines from the Clapeyron equation ($dP/dT_m = \Delta H/T_m \Delta V$), where the specific volume (ΔV) and enthalpy (ΔH) are obtained from one-phase simulations, whereas the melting temperature (T_m) at which these quantities are computed is taken from the two-phase results (Fig. 1).

A detailed error analysis shows that at high-pressure (>500 GPa) errors introduced by finite size effects and by neglecting the finite electronic temperature (T_e) are of opposite sign. The melting temperature of diamond increases by ≈ 200 K in going from 128- to 256-atom simulation cells (at $P = 1,000$ GPa), but it decreases by ≈ 500 K when considering $T_e \approx 8,000$ K in our calculation with

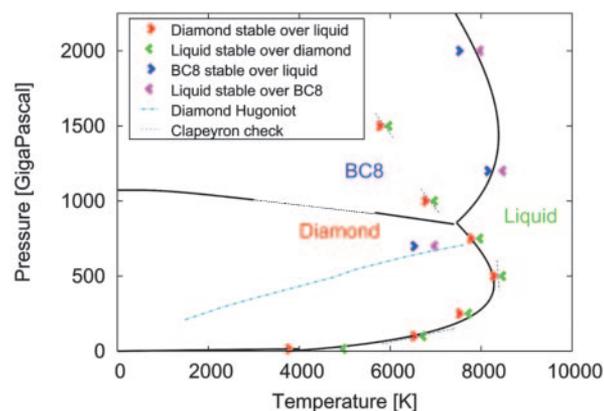


Fig. 1. Calculated phase diagram of carbon at high pressure omitting the graphite stability region for clarity. Points indicate the final two-phase simulation data used to bracket the melting temperatures. The plotted melting lines are a fit to the melting points by using the three-parameter Kechin melting equation (34). For a consistency check, the slopes of the melting curves (marked with short dotted segments) were computed independently from the Clapeyron equation (35). The phase boundary between diamond and BC8 is calculated within the quasiharmonic approximation at low temperature and from the three-phase Clapeyron identity Eq. 1 near the triple point. The triple point is located at $T = 7,445$ K and $P = 850$ GPa; metastable extensions of melting lines of BC8 and diamond are calculated and presented for illustrative purposes. The diamond Hugoniot (dashed line) is shown for comparison with experimental results (see text).

128-atom simulation cells (at 500–750 GPa and 8,000 K). Introducing electronic temperature likely reduces the directional character of the interatomic forces in the liquid, thus increasing the number of effective degrees of freedom of the ions, which, in turn, results in an increase of the ionic entropy and, hence, a reduction of the free energy of the liquid relative to that of the solid. In contrast, the introduction of an electronic temperature has little or even a null effect on the insulating solid. In addition, we found a decrease of the melting temperature of ≈ 200 K at 1,000 GPa on decreasing the pseudopotential cutoff radius; this effect diminishes rapidly at lower pressures.

Although we estimate our numerical errors to be approximately $\pm 150/250$ K at fixed pressure, the error introduced by DFT is more difficult to quantify. We note that in cases for which the electronic charge density tends to be more homogeneous in the liquid phase than in the solid, the generalized gradient-corrected approximation (GGA) is expected to favor the former phase and, thus, to underestimate the melting temperature, which is the case, for example, for Al, Si, and LiH, and it is likely to be so for diamond and BC8. Although quantitative estimates for carbon are difficult to make in the absence of experimental data and calculations beyond the GGA, it is worth noting that, for Si, the melting temperature at ambient conditions is underestimated by $\approx 10\%$, and this error is expected to decrease as pressure is increased (31, 36).

The melting line of diamond obtained in our calculations is consistent with that reported by Grumbach and Martin (21), although the triple point found here is at a lower pressure than one might infer from their results. Our predicted diamond melting temperatures are also lower than those obtained in recent calculations with semiempirical potentials (26), especially in the high-pressure region. The reason for this discrepancy may come from the data set to which the empirical potentials have been fitted, which only contains low pressure, mostly tetrahedral, structures. We note that the gradual change in coordination number found here is the key for understanding the changes in the volume of the liquid relative to the solid as pressure is increased, and it is responsible for the existence of a maximum on the melting line, as we discuss below.

As apparent from Fig. 1, both diamond and BC8 melting lines

^{||}The discrepancy with previous reports on the stability of R8 stems from the numerical accuracy of the calculations that have appeared in the literature, which were not fully converged as a function of plane wave cutoff and k -point sampling.

We thank Francois Gygi, Tadashi Ogitsu, Eric Schwegler, Xiaofei Wang, and Roberto Car for many useful discussions and Roger Falcone for constant help and support and for a critical reading of the manuscript. This work was performed under the auspices of the U.S.

Department of Energy at the University of California/Lawrence Livermore National Laboratory under Contract W-7405-Eng-48. S.A.B. was supported by the Natural Sciences and Engineering Research Council of Canada.

1. Lu, P. J., Yao, N., So, J. F., Harlow, G. E., Lu, J. F., Wang, G. F. & Chaikin, P. M. (2005) *Archaeometry* **47**, 1–12.
2. Bundy, F., Hall, T., Strong, H. M. & Wentorf, R. H. (1955) *Nature* **176**, 51–54.
3. Bundy, F. P., Bassett, W. A., Weathers, M. S., Hemley, R. J., Mao, H. K. & Goncharov, A. F. (1996) *Carbon* **34**, 141–153.
4. Ross, M. (1981) *Nature* **292**, 435–436.
5. Hubbard, W. B. (1981) *Science* **214**, 145–149.
6. Benedetti, L. R., Nguyen, J. H., Caldwell, W. A., Liu, H., Kruger, M. & Jeanloz, R. (1999) *Science* **286**, 100–102.
7. Segretain, L., Chabrier, G., Hernanz, M., Garcia-Berro, E., Isern, J. & Mochkovitch, R. (1994) *Astrophys. J.* **434**, 641–651.
8. Metcalfe, T. S., Montgomery, M. H. & Kannan, A. (2004) *Astrophys. J.* **605**, L133–L136.
9. Kuchner, M. J. & Seager, S. (2005) *Astrophys. J.*, in press.
10. Yin, M. T. & Cohen, M. L. (1984) *Phys. Rev. Lett.* **50**, 2006–2009.
11. Yin, M. T. (1984) *Phys. Rev. B Condens. Matter* **30**, 1773–1776.
12. Hu, J. Z., Merkle, L. D., Menoni, C. S. & Spain, I. L. (1994) *Phys. Rev. B Condens. Matter* **34**, 4679–4684.
13. Nelmes, R. J., McMahon, M. I., Wright, N. G., Allan, D. R. & Loveday, J. S. (1993) *Phys. Rev. B Condens. Matter* **48**, 9883–9886.
14. Occelli, F., Loubeyre, P. & Letoullec, R. (2003) *Nat. Mater.* **2**, 151–154.
15. Bradley, D. K., Eggert, J. H., Hicks, D. G., Celliers, P. M., Moon, S. J., Cauble, R. C. & Collins, G. W. (2004) *Phys. Rev. Lett.* **93**, 195506–195509.
16. Weathers, M. S. & Bassett, W. A. (1987) *Phys. Chem. Miner.* **15**, 105–112.
17. Togaya, M. (1990) *High Press. Res.* **4**, 342–348.
18. Galli, G., Martin, R. M., Car, R. & Parrinello, M. (1989) *Phys. Rev. Lett.* **63**, 988–991.
19. Galli, G., Martin, R. M., Car, R. & Parrinello, M. (1990) *Science* **250**, 1547–1549.
20. Wu, C. J., Glosli, J. N., Galli, G. & Ree, F. H. (2002) *Phys. Rev. Lett.* **89**, 135701–135704.
21. Grumbach, M. P. & Martin, R. (1996) *Phys. Rev. B* **54**, 15730–15741.
22. Brenner, D. W. (1990) *Phys. Rev. B* **42**, 9458–9471.
23. van Thiel, M. & Ree, F. H. (1992) *High Press. Res.* **10**, 607–610.
24. Glosli, J. N. & Ree, F. H. (1999) *Phys. Rev. Lett.* **82**, 4659–4662.
25. Fried, L. E. & Howard, W. M. (2000) *Phys. Rev. B* **61**, 8734–8743.
26. Ghiringhelli, L. M., Los, J. H., Meijer, E. J., Fasolino, A. & Frenkel, D. (2005) *Phys. Rev. Lett.* **94**, 145701–145704.
27. Clark, S. J., Ackland, G. J. & Crain, J. (1995) *Phys. Rev. B Condens. Matter* **52**, 15035–15038.
28. Biswas, R., Martin, R. M., Needs, R. J. & Nielsen, O. H. (1987) *Phys. Rev. B Condens. Matter* **35**, 9559–9568.
29. Bundy, F. P. & Kasper, J. S. (1967) *J. Chem. Phys.* **46**, 3437–3446.
30. Ogitsu, T., Schwegler, E., Gygi, F. & Galli, G. (2003) *Phys. Rev. Lett.* **91**, 175502–175505.
31. Alfe, D. (2003) *Phys. Rev. B Condens. Matter* **68**, 064423–064426.
32. Bonev, S. A., Schwegler, E., Ogitsu, T. & Galli, G. (2004) *Nature* **431**, 669–672.
33. Steinhardt, P. J., Nelson, D. R. & Ronchetti, M. (1983) *Phys. Rev. B Condens. Matter* **28**, 784–805.
34. Kechin, V. V. (2002) *Phys. Rev. B Condens. Matter* **65**, 052102–052105.
35. Zemansky, M. W. (1968) *Heat and Thermodynamics* (McGraw-Hill, New York), 5th Ed.
36. Sugino, O. & Car, R. (1995) *Phys. Rev. Lett.* **74**, 1823–1826.
37. Xie, J., Chen, S. P., Tse, J. S., de Gironcoli, S. & Baroni, S. (1999) *Phys. Rev. B Condens. Matter* **60**, 9444–9449.
38. Crain, J., Clark, S. J., Ackland, G. J., Payne, M. C., Milman, V., Hatton, P. D. & Reid, B. J. (1994) *Phys. Rev. B Condens. Matter* **49**, 5329–5340.
39. Tsuchiya, J., Tsuchiya, T. & Wentzcovitch, R. M. (2005) *J. Geophys. Res.* **110**, B02204–B02209.
40. Schulz, M. & Weiss, H. (1982) *Landolt-Börnstein Tables, Numerical Data and Functional Relationships in Science and Technology*, ed. Madelung, O. (Springer, Berlin), Vol. New Series 17a.
41. Dean, P. J., Lightowers, E. C. & Wight, D. R. (1965) *Phys. Rev.* **140**, A352–A368.
42. Surh, M. P., Louie, S. G. & Cohen, M. (1992) *Phys. Rev. B Condens. Matter* **45**, 8239–8247.
43. Kubo, R. (1957) *J. Phys. Soc. Jpn.* **12**, 570–586.
44. Papadopoulos, A. D. & Anastassakis, E. (1990) *Phys. Rev. B Condens. Matter* **43**, 9916–9923.
45. Souza, I., Martin, R. M., Marzari, N., Zhao, X. & Vanderbilt, D. (2000) *Phys. Rev. B Condens. Matter* **62**, 15505–15520.
46. Stevenson, D. J. (1982) *Annu. Rev. Earth Planet. Sci.* **10**, 257–295.
47. Guillot, T. (1999) *Science* **286**, 72–77.
48. Guillot, T. (2005) *Annu. Rev. Earth Planet. Sci.* **33**, 493–530.
49. Troullier, N. & Martins, J. L. (1991) *Phys. Rev. B Condens. Matter* **43**, 1993.